

# Dynamic Compression of Iron Carbide at Planetary Core Conditions

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Carnegie Institution for Science

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# Our Team

**Joint experimental & theory project**

## **Experiments:**

Sally Tracy, & **Postdoctoral Fellow** (Carnegie)

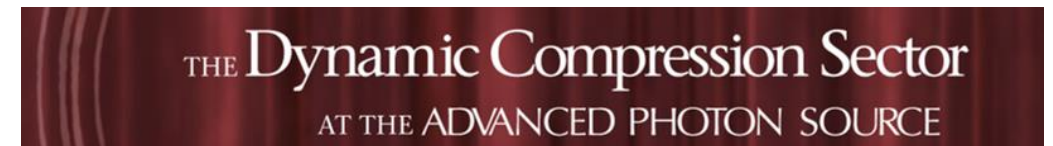
## **Simulations:**

Ivan Oleynik, Kien Nguyen Cong (USF)

## **National Lab Collaborators:**

Ray Smith, Federica Coppari, Jon Eggert, Marius Millot, Chris McGuire, & Samantha Clarke (LLNL)

Support from User Facility Staff at DCS and LLE



# Planetary Cores

**Structure & properties of core materials** at the P-T conditions of Earth's core and the core's of larger super-Earth exoplanets

## Earth's Core:

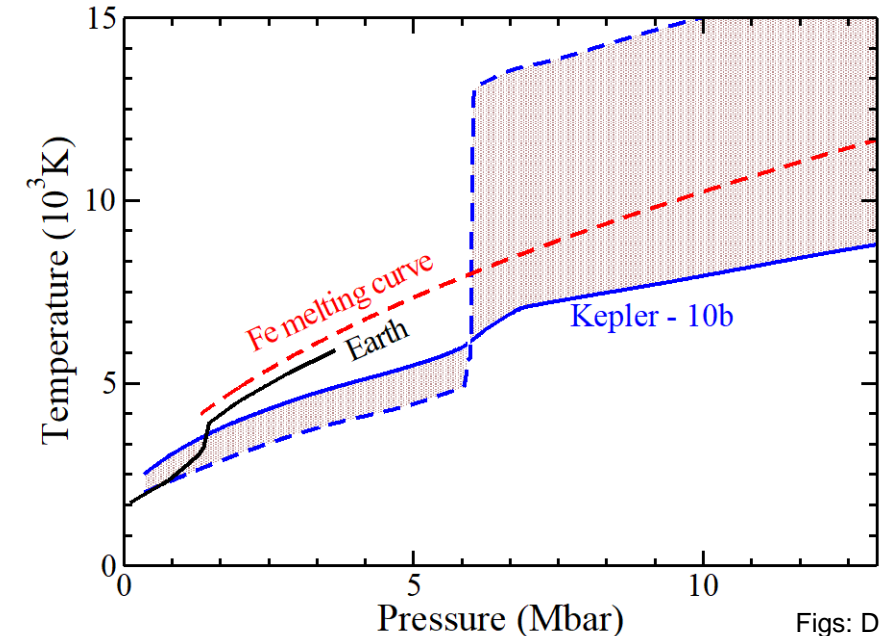
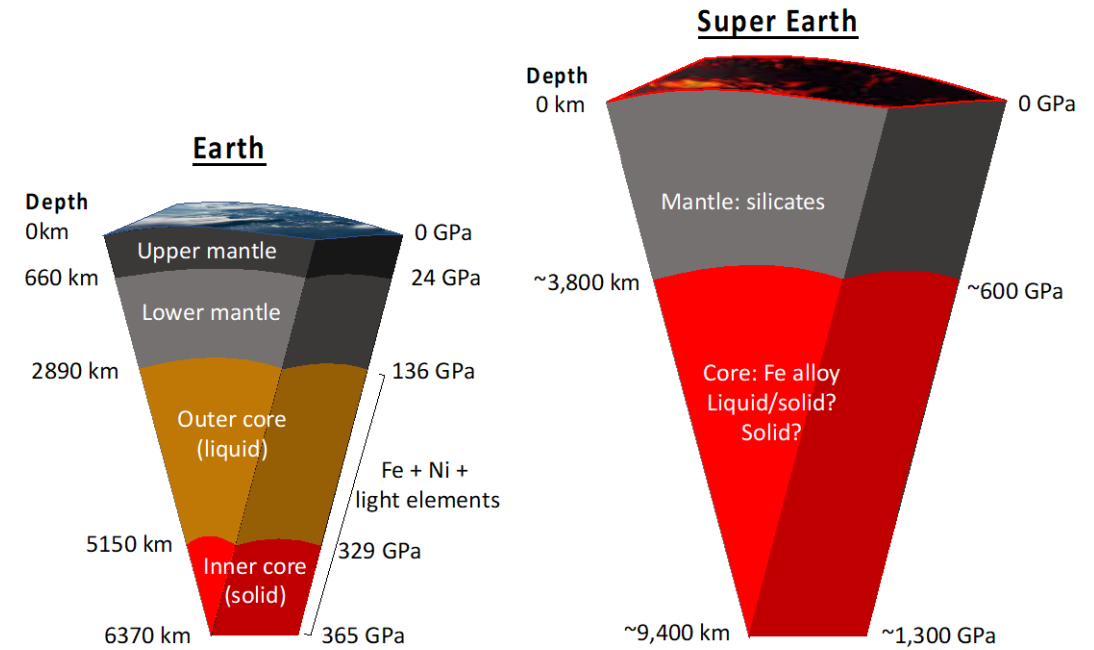
Liquid outer core & crystalline inner core  
Fe-Ni alloy with 10% lighter elements (C, Si, O...)

## Super-Earths:

Abundant population of Earth-like planets with up to 10-Earth masses and internal Ps up to 4 TPa

Super-Earth are also expected to incorporate light elements

Focus of this project is on **carbon** an important candidate light element for Earth's core as well as C-rich super-Earths



Figs: Duffy & Smith 2019



# Effects of Light Elements

## Light elements effect melting temperature

Relative slopes of the liquidus and the planetary adiabat will determine if a planet has a coexisting liquid outer and solid inner core

Liquid outer core allows for a dynamo to form, generating a magnetic field which plays a role in habitability

## Complex structures in the Earth's solid inner core

- Thin outermost layer
- West & East hemispheres with different elastic properties
- Distinct innermost inner core

Source of heterogeneities may be related to light-element-rich alloy phases.

Changes in crystallization over time that effect concentration of light elements in the precipitating solids.

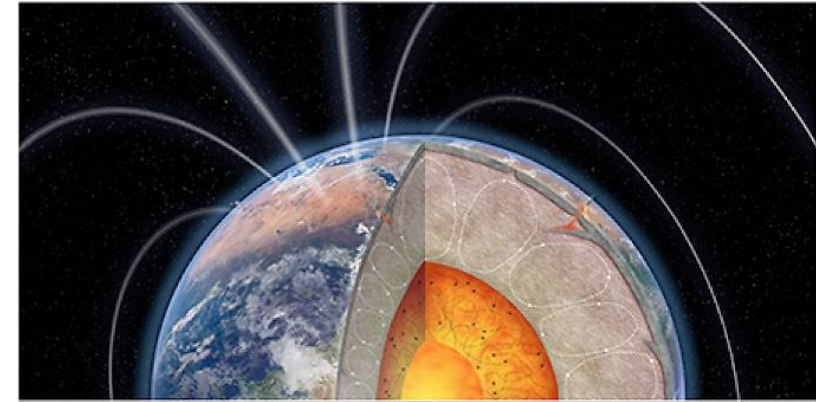


Fig: UC Santa Cruz

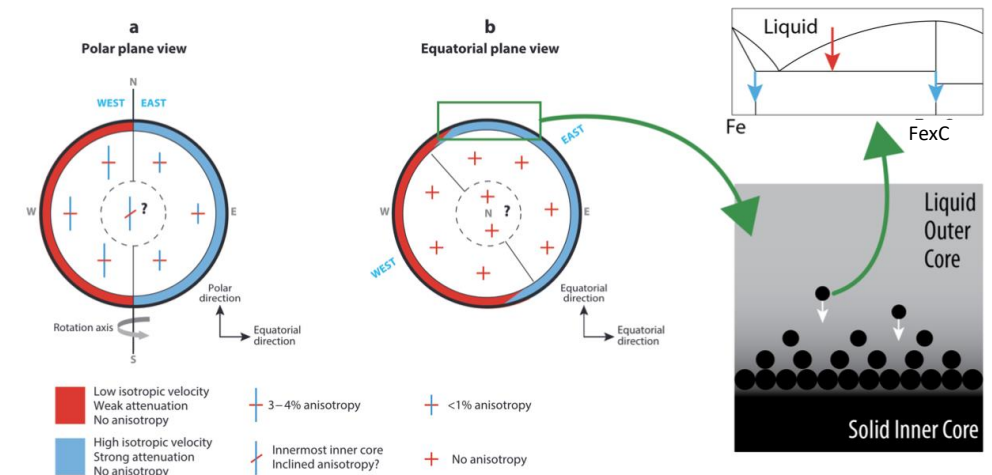


Fig: Deuss et al.

# Iron Carbide Phase Diagram

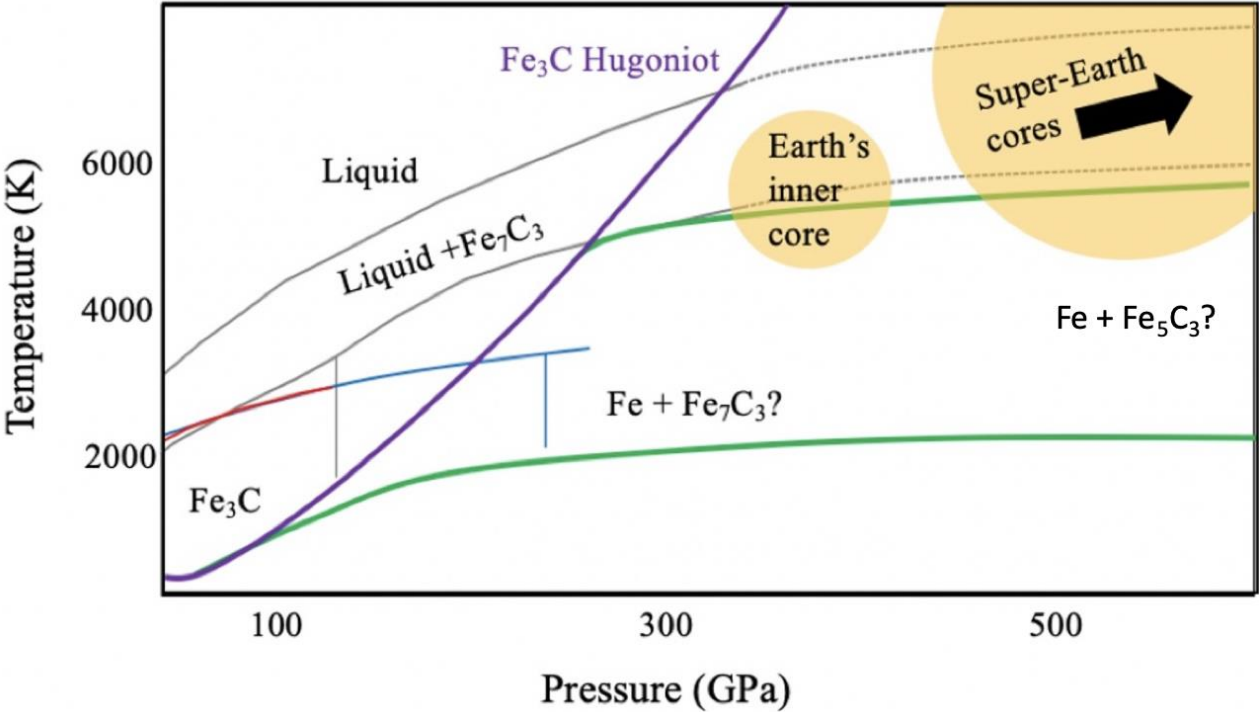
Understanding of C incorporation into the core is limited by poor constraints on the high-P Fe-C phase diagram

Carbon has limited solubility in HCP-Fe

Solid cores likely include Fe-carbide phases

Models based on long extrapolations predict  $\text{Fe}_3\text{C}$  was 1<sup>st</sup> phase to crystallize during solidification of the Earth's core

Recent results suggest  $\text{Fe}_7\text{C}_3$  may instead be the favored phase such that  $\text{Fe}_3\text{C}$  melts incongruently



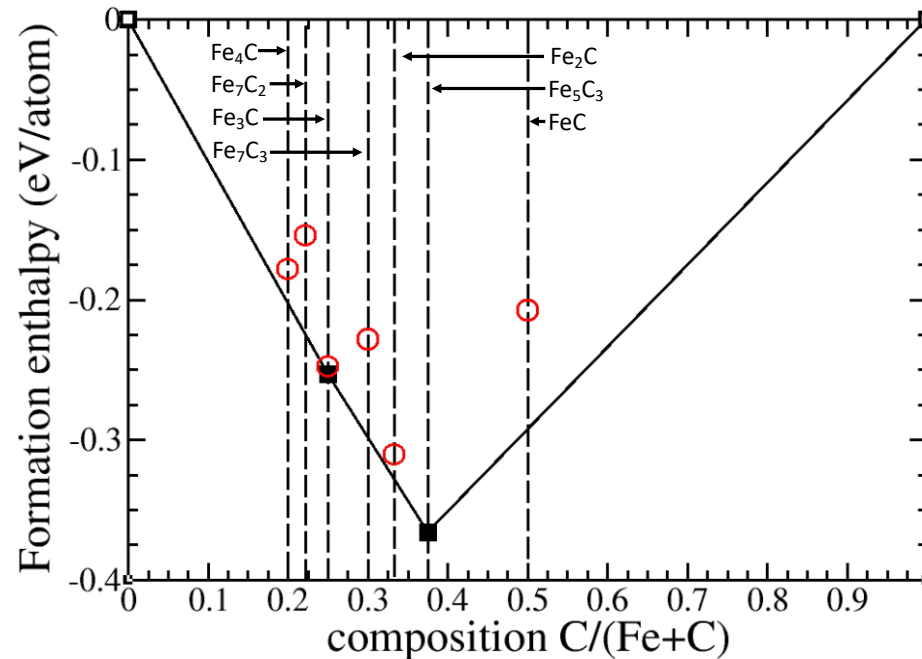
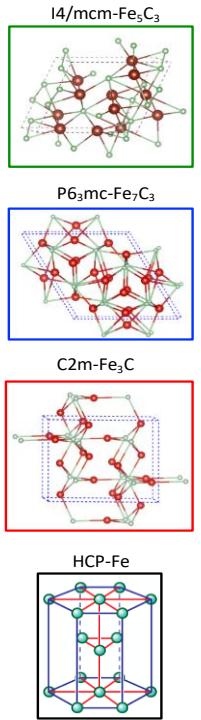
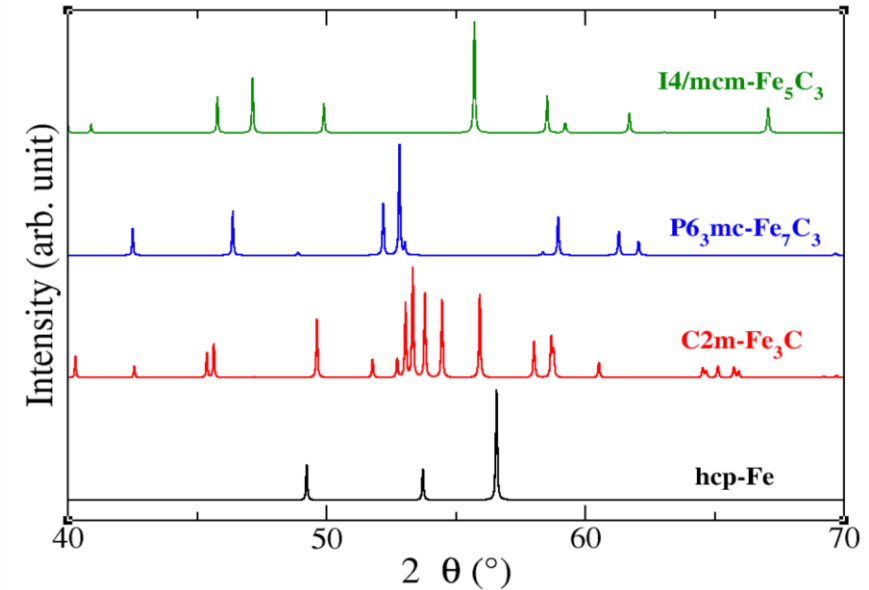
No experimental data at P-T conditions of inner core or cores of larger super-Earths

# Structure Search Calculations

Theoretical calculations predict  $\text{Fe}_2\text{C}$  to be more stable than either  $\text{Fe}_3\text{C}$  or  $\text{Fe}_7\text{C}_3$  at inner core conditions

Our recent structure search results indicate a new  $\text{Fe}_5\text{C}_3$  structure is the lowest enthalpy phase at 500 GPa

Based on these results, we expect  $\text{Fe}_3\text{C}$  may break down into HCP Fe + carbide phases of different stoichiometries



Kien Nguyen Cong (USF)



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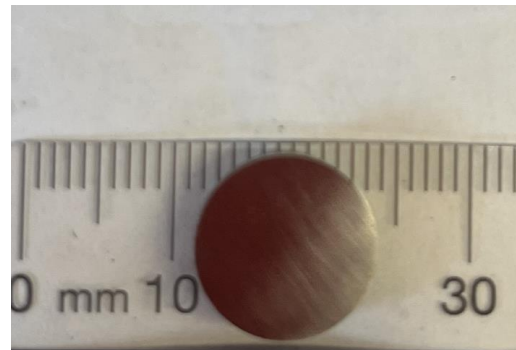


# Fe<sub>3</sub>C Synthesis at Carnegie

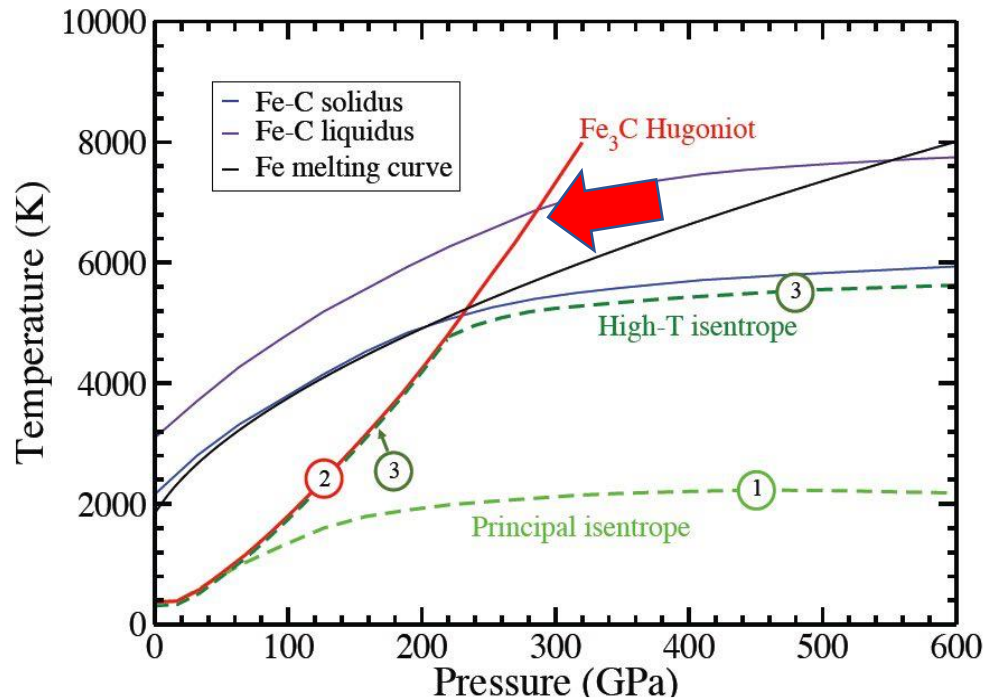
Start with Fe<sub>3</sub>C allowing us to explore Fe-C phase stability with a well-constrained stoichiometry (6.7 wt% C)

98% dense polycrystalline Fe<sub>3</sub>C sample synthesized by solid phase reaction between Fe and graphite powder at high-PT

Sintered at 4.8 GPa and 1100 °C for 15 min, using a large-volume cubic press



# Structure of Fe<sub>3</sub>C Under Shock Loading up to 250 GPa



X-ray Diffraction (XRD) along the **Fe<sub>3</sub>C Hugoniot**

## Questions

What is the crystal structure of Fe<sub>3</sub>C under shock compression?

Can we constrain the Fe-C melt curve above 100 GPa?

Melting → Loss of XRD peaks & appearance of diffuse scattering

Is melting is congruent or incongruent?

## Implications

Explore stabilization of carbide phases in Earth's inner core

Assess carbon's effect on suppression of the liquidus at Earth-core conditions





# In Situ Diffraction Under Laser Shock Loading at DCS



DCS

Laser ( $<80$  J, 5-10 ns pulse length) is focused on a plastic ablation layer generating a plasma which expands backwards sending a shock wave into the sample

During the shock state, sample is probed with XRD data is recorded using a transmission geometry

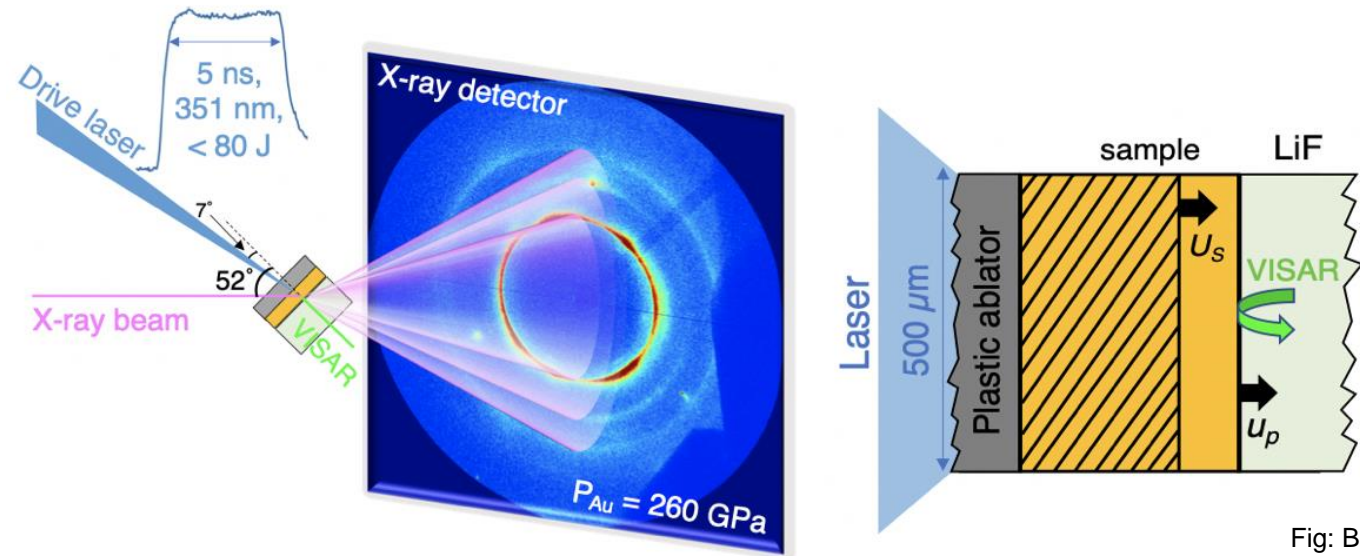


Fig: Briggs et al.

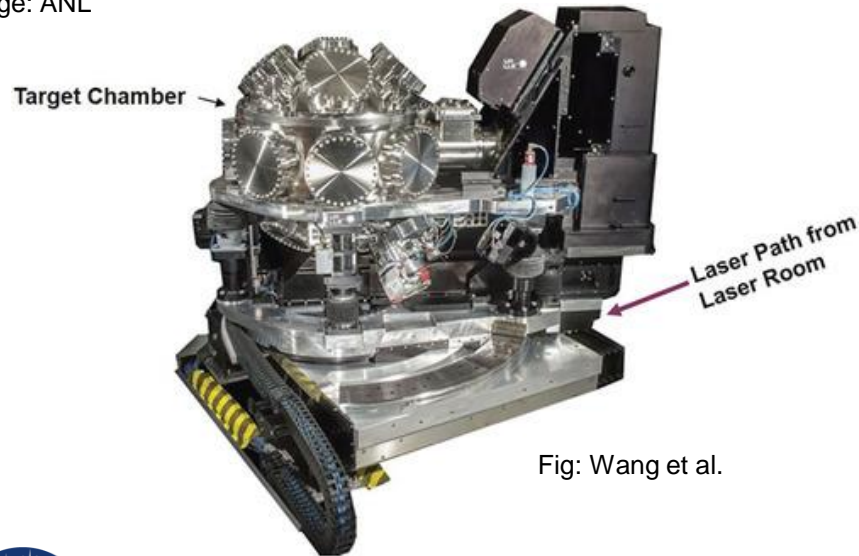
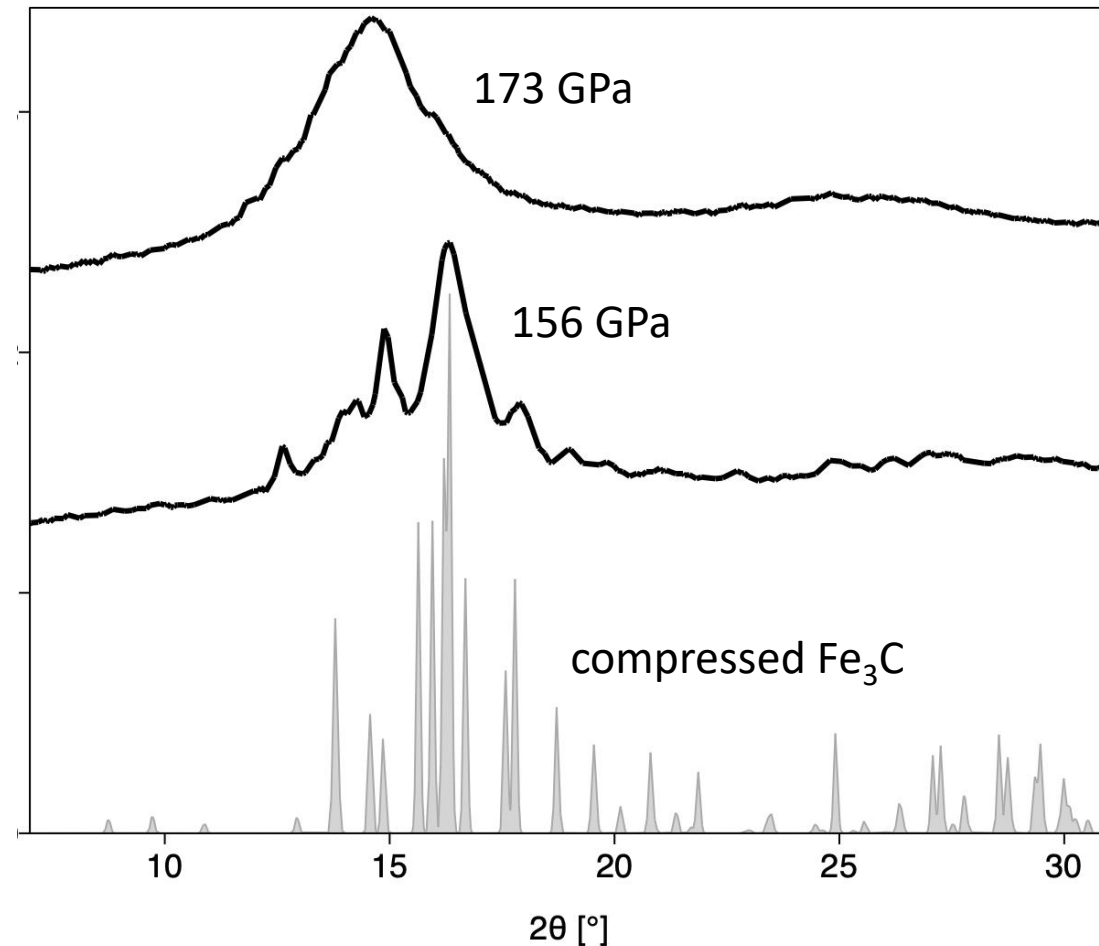
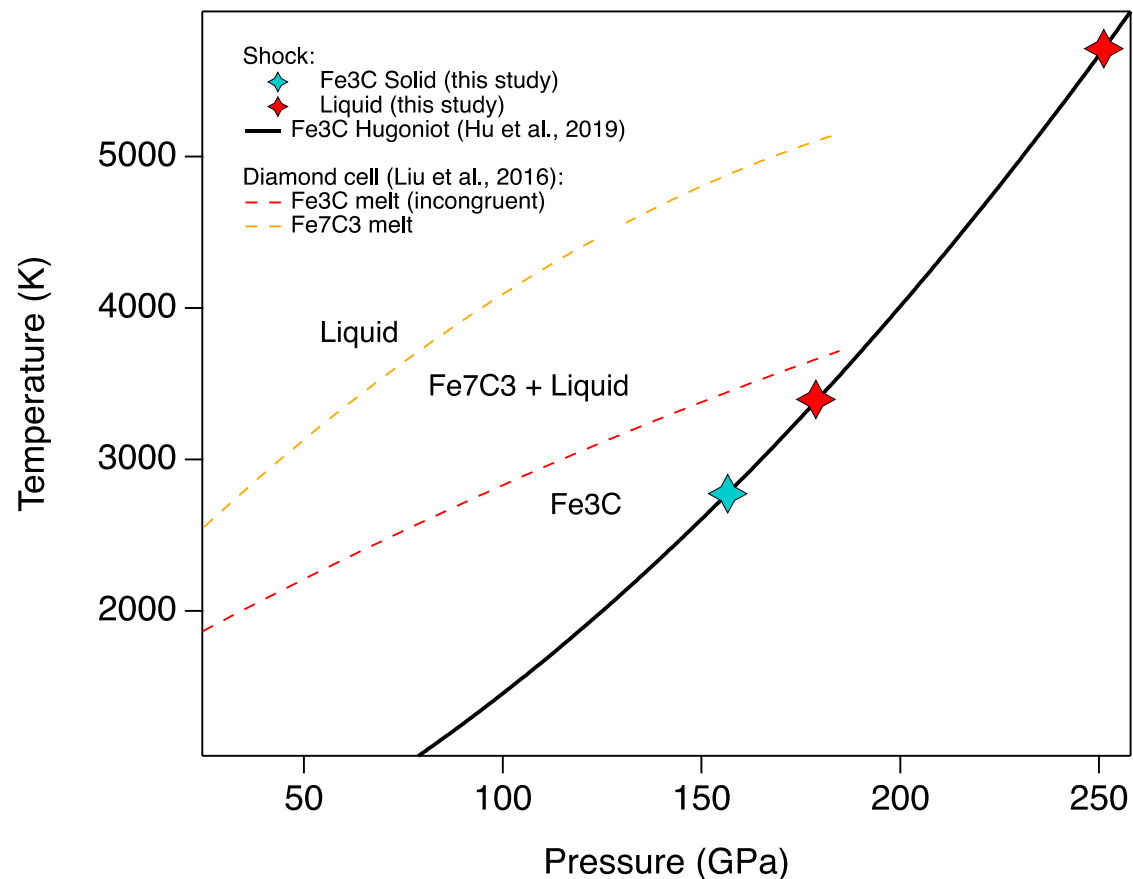


Fig: Wang et al.



# Preliminary DCS Results



**Fe<sub>3</sub>C melts at 160 GPa into a single liquid along the Hugoniot**

Contrast expected incongruent melting (Fe<sub>3</sub>C  $\rightarrow$  Fe<sub>7</sub>C<sub>3</sub> + Lq.)

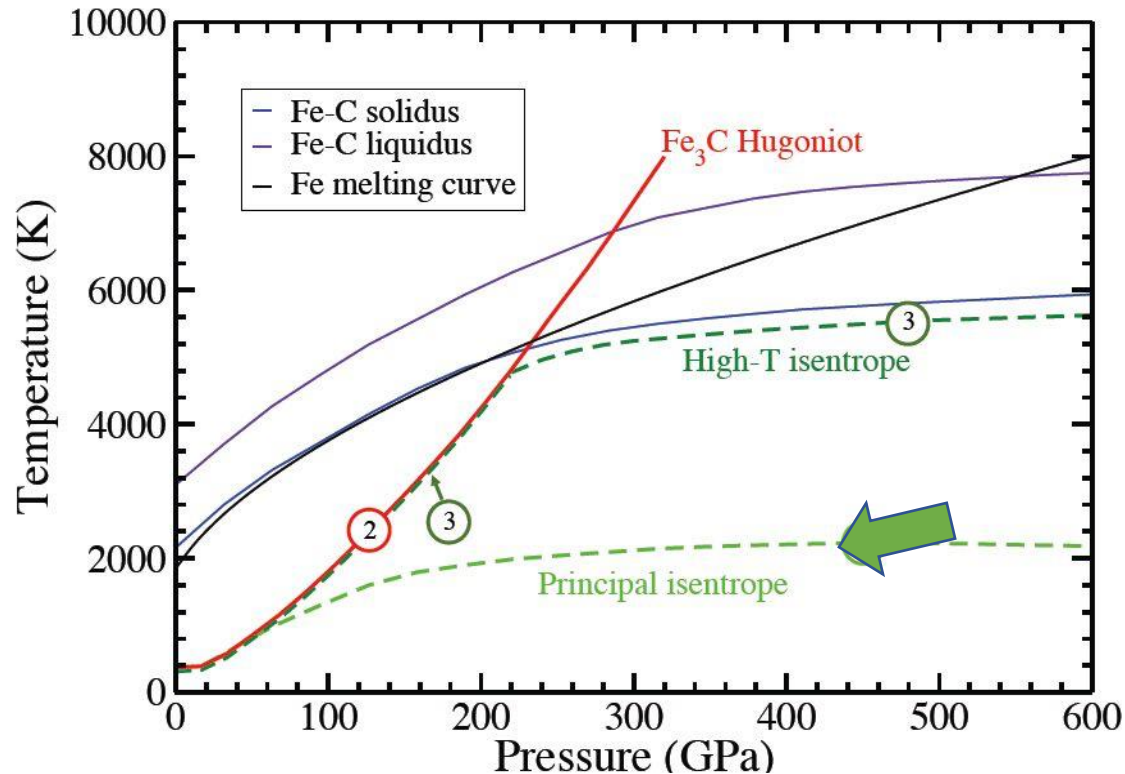
**Future:** Quantitative analysis of liquid structures between 160-250 GPa



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# Structure of $\text{Fe}_3\text{C}$ Under Ramp Loading up to 600 GPa



## Questions

What is structure and EOS of iron carbide at conditions of the super-Earth cores?

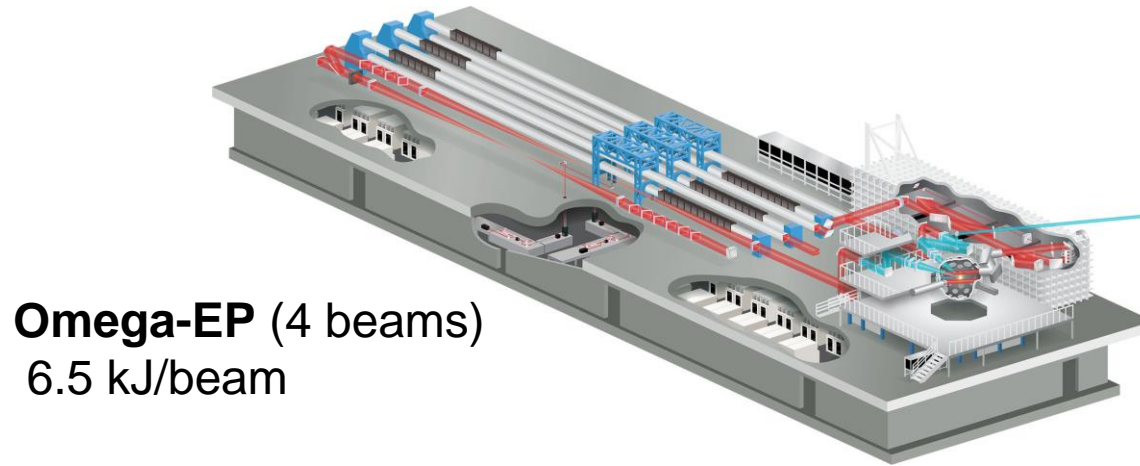
Does  $\text{Fe}_3\text{C}$  undergoes decomposition into HCP Fe and higher carbon content carbides (e.g.,  $\text{Fe}_3\text{C}_5$ ) at ultrahigh pressures?

## Implications

Explore stabilization of carbide phases in cores of up to 2-Earth mass terrestrial exoplanets

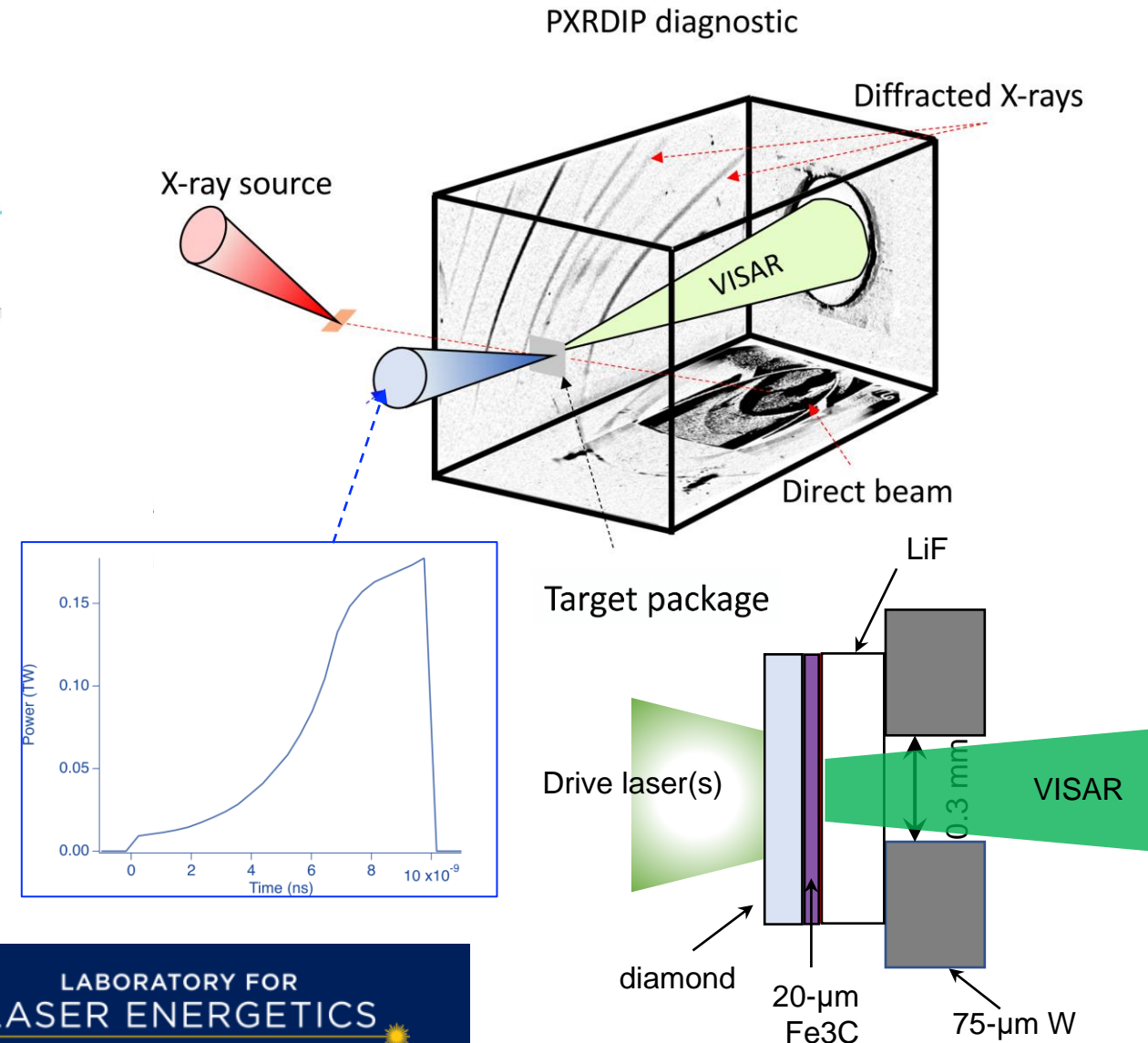


# Plasma X-Ray Source at Omega EP Laser



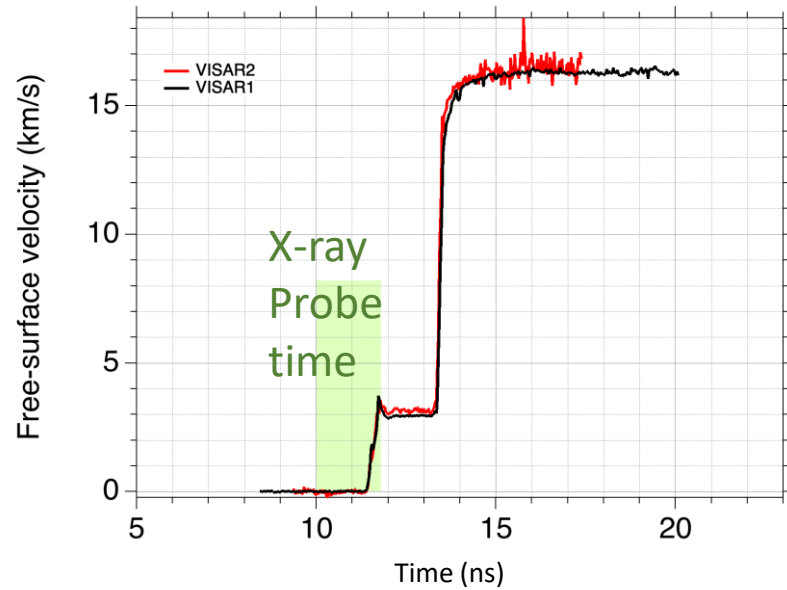
High-power multibeam laser facility where we can drive samples with a kJ pulse & tailored pulse shapes

Drive a thin Fe foil with a 2<sup>nd</sup> high-intensity laser pulse which induces  $\text{He}_\alpha$  emission of quasimonochromatic X-rays





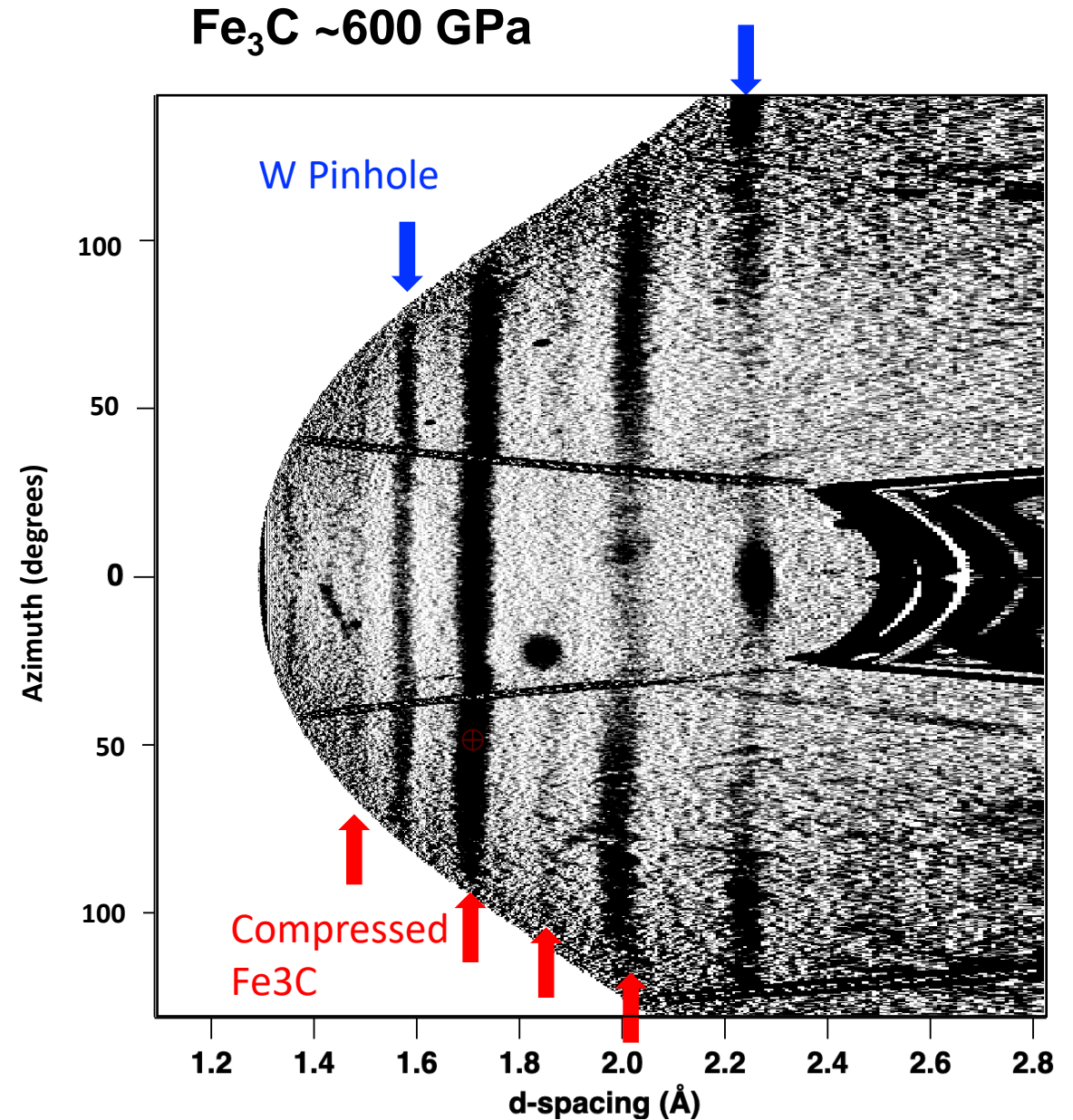
# Preliminary EP Results



Collected 7 shots between 200-600 GPa

**Orthorhombic  $\text{Fe}_3\text{C}$  phase observed up to highest pressure**

**No evidence for phase transition or breakdown into HCP-Fe + carbides**





# Up Next: Recrystallization at Omega EP

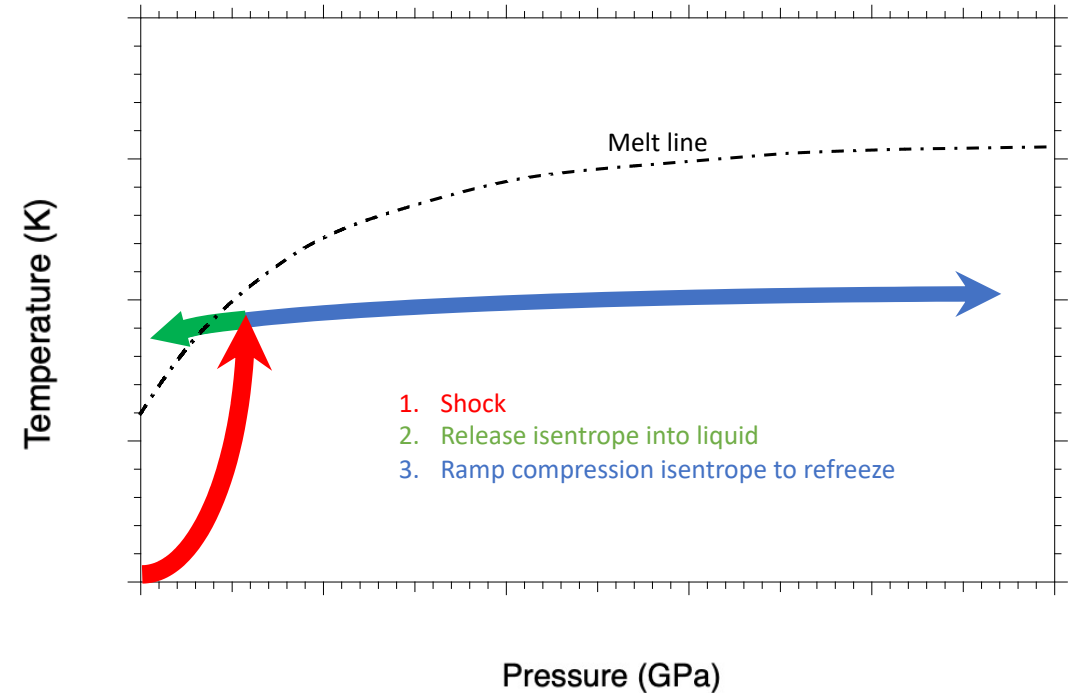
## What role does kinetics have on our observations?

$\text{Fe}_3\text{C}$  predicted to break down into HCP-Fe + carbide phases of different stoichiometries

Decomposition reactions are diffusion limited so may not be observed in ns dynamic compression experiments

## Shock-Ramp – Melt-Refreeze

Recrystallization of the stable phase assemblage from the liquid



## Implications

Direct measure of phase(s) crystallizing from the high-P melt at the pressures of Earth's inner-outer core boundary

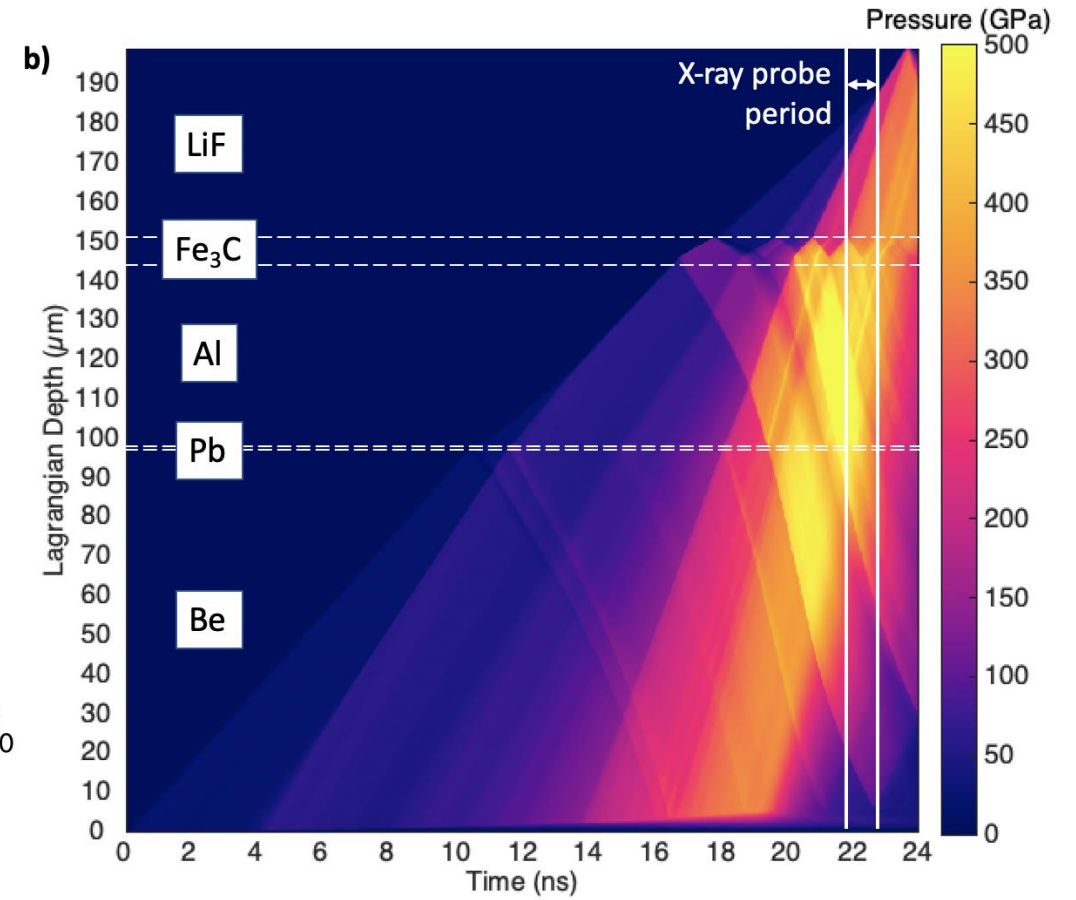
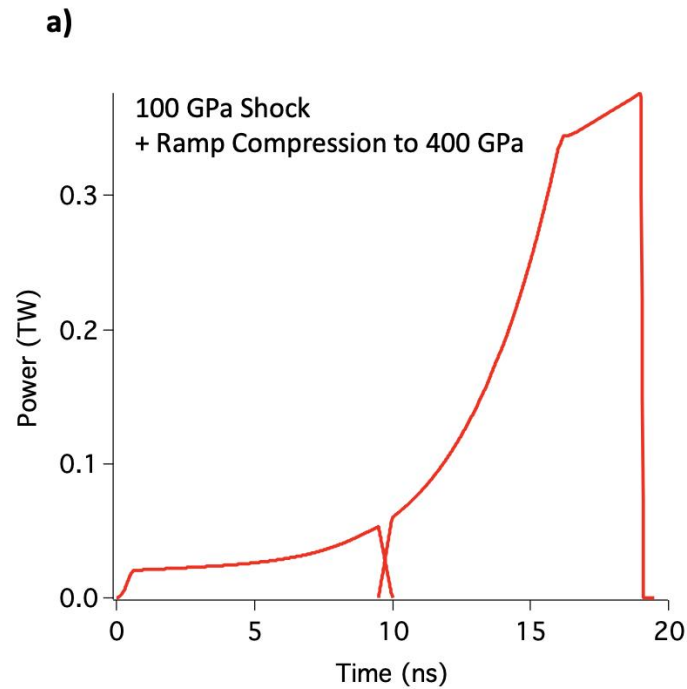
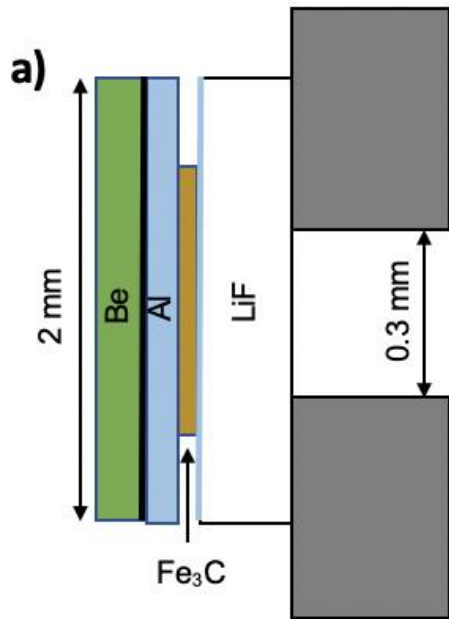
Understand kinetics of nanosecond dynamic experiments



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# Shock-Ramp experiments



# Conclusions

- Fe-C structure searches show evidence for stability of  $\text{Fe}_7\text{C}_3$  and  $\text{Fe}_5\text{C}_3$  at planetary core pressures.
- $\text{Fe}_3\text{C}$  melts at 160 GPa into a single liquid along the Hugoniot
- Orthorhombic  $\text{Fe}_3\text{C}$  was observed under ramp compression up to 600 GPa, with no evidence for phase transition or breakdown into HCP iron and C-rich carbides

# Future Directions

- Shock-ramp and melt-refreeze experiments at Omega EP
- Liquid structure analysis of DCS Data
- Molecular Dynamics calculations of temperatures along variable compression pathways



# Shock-Ramp experiments 400 GPa

